Radio as a Unique Probe To Access the Largest Cosmic Scales

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Primordial Non-Gaussianity

$$\Phi = \phi + f_{\rm NL} * (\phi^2 + \langle \phi^2 \rangle)$$

Predicted in many scenarios of inflation

Primordial Non-Gaussianity

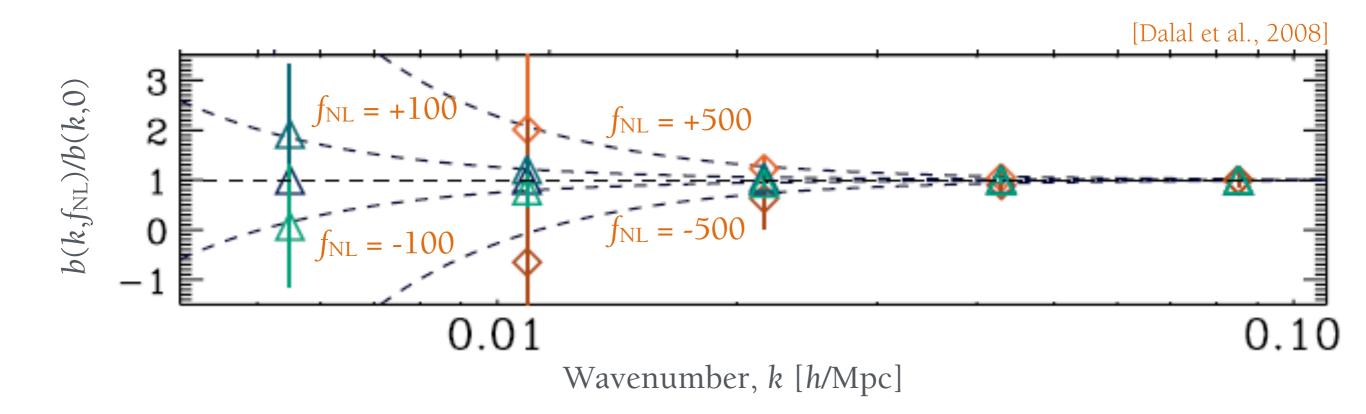
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- Predicted in many scenarios of inflation
- Tightest available constraints from CMB: $|f_{NL}| < 10$ [Planck Collaboration, 2015]

Primordial Non-Gaussianity

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• Accuracy of O(1) possible w/ future large-scale galaxy surveys









Accessing the Largest Scales

- Forecasts on primordial non-Gaussianity for galaxy surveys
 - $\sigma(f_{\rm NL}) \sim 1.5$ [SC, Santos & Maartens, 2015]

• $\sigma(f_{NL}) \sim 2 - 4$ [Carbone, Matarrese & Verde, 2008; Giannantonio et al., 2012]







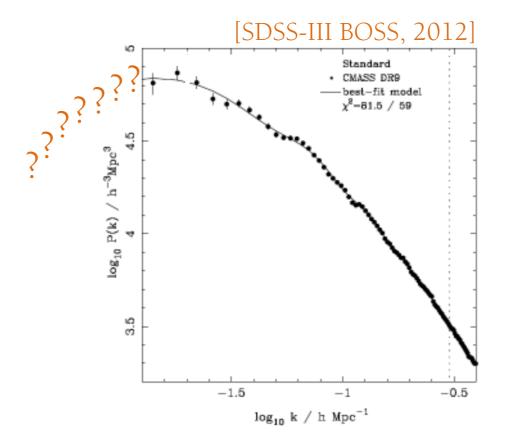
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Intensity Mapping

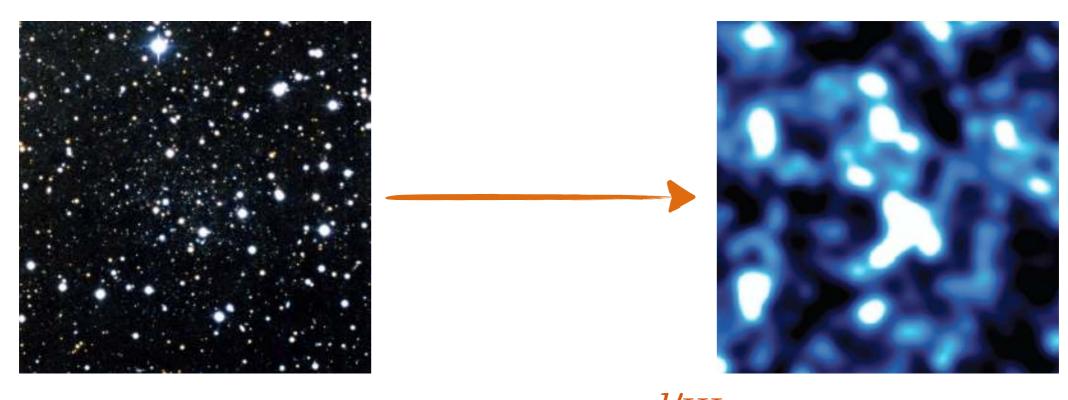


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- Developed for radio dish surveys / interferometers: SKA $\sigma(f_{NL}) \sim 1$ [SC, Santos, Ferreira & Ferramacho, PRL 2013]



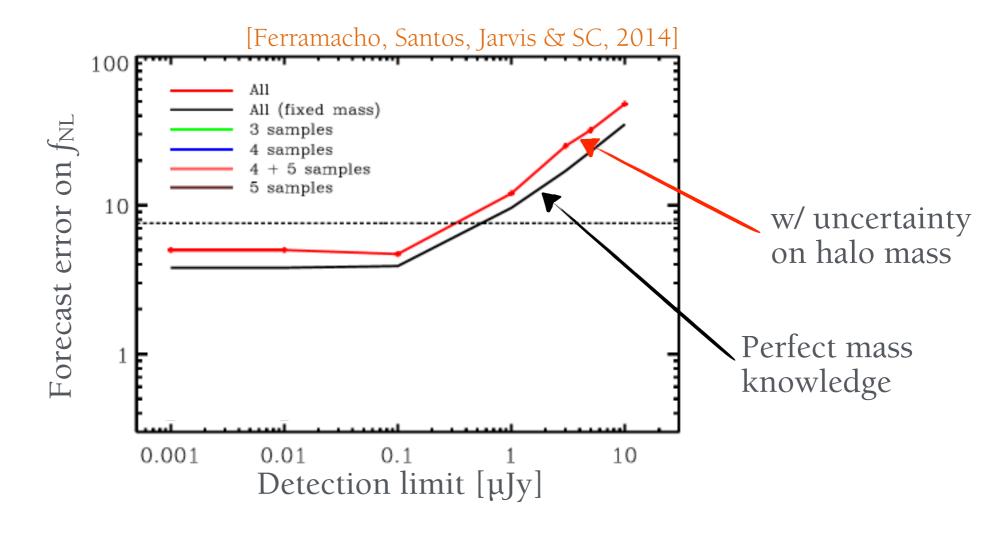
Multi-Tracer Technique

• Comparing the relative clustering of different populations of tracers [Seljak, 2009]



Multi-Tracer Technique

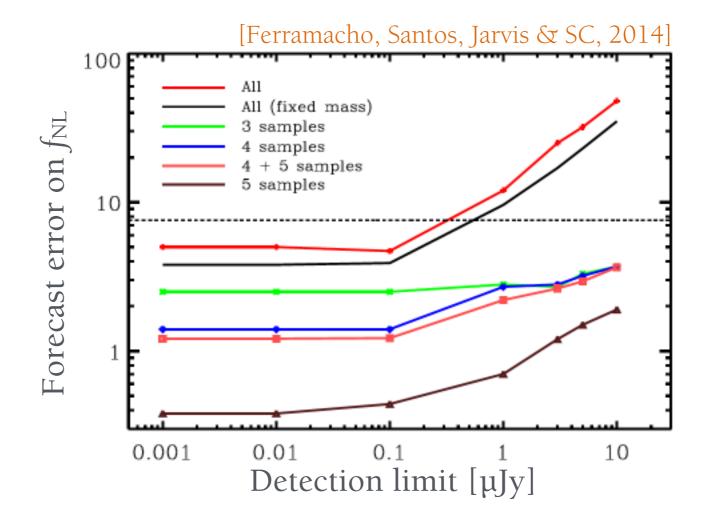
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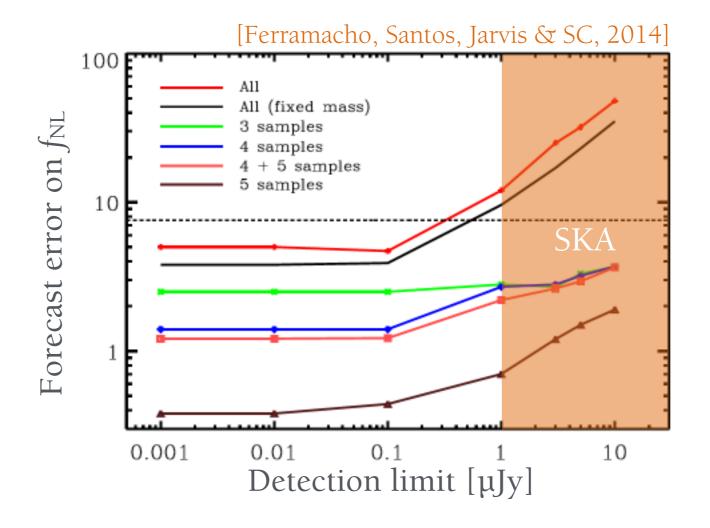
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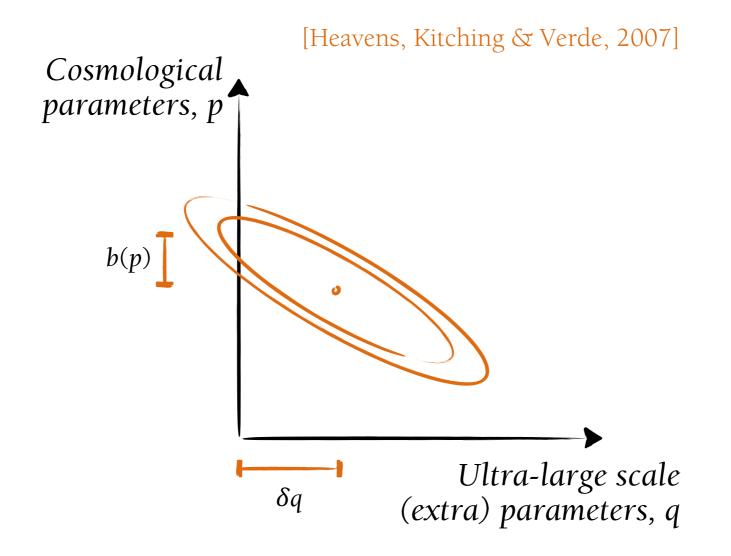
Sumary

- Constraints on the properties of density perturbations on extremely large scales will improve our understanding of the law of gravity
 - The largest scales are a playground for possibly unknown physics (modified gravity?)
 - They can further confirm Einstein's general relativity (tests for relativistic corrections)
- They can also improve our knowledge of inflation and early Universe (e.g. primordial non-Gaussianity)
- Radio has a unique potential in accessing the largest cosmic scales! (HI intensity mapping, multi-tracer technique...)



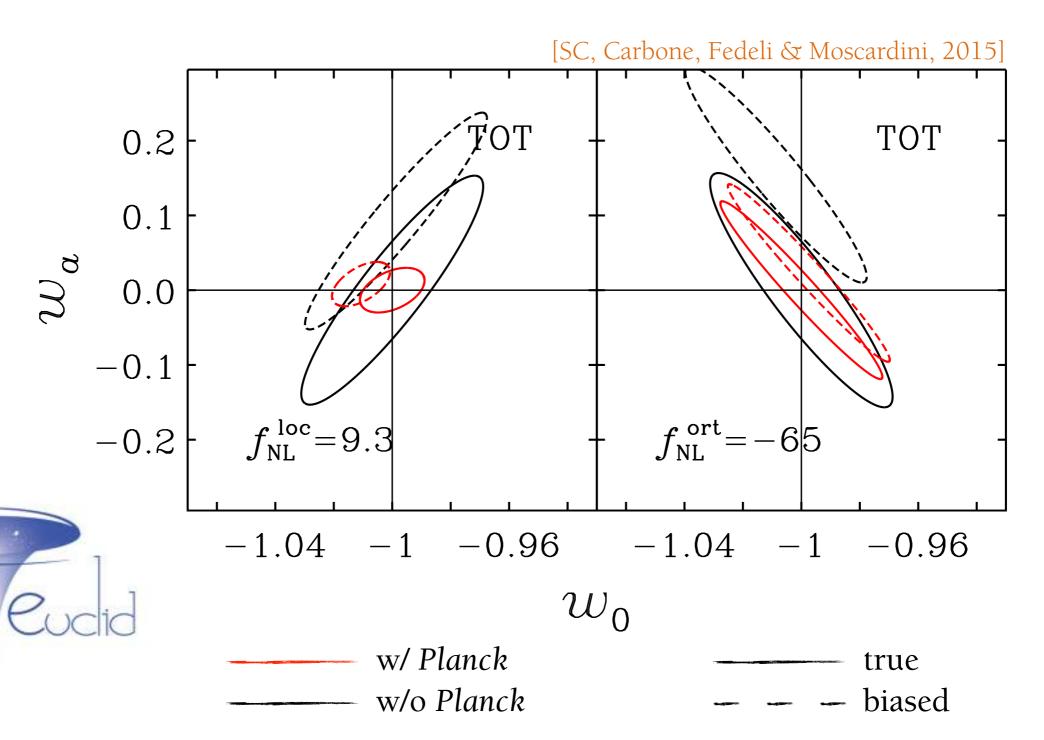
Ultra-Large Scales Matter!

 Neglecting ultra-large scale effects can undermine future cosmological experiments' accuracy





Neglecting non-Gaussianity





Neglecting Relativistic Effects

